

Introduction to astrophysical plasma physics:

(1)

Relevant scales and equations

based on R. Kulsrud „Plasma Physics for Astrophysics“

- What is plasma? What is the most general way of describing the plasma?
 - Distribution function
 - Kinetic equation, Maxwell's equations
 - Collisions
- Particle motion in uniform fields
 - Drifts
 - Larmor radius
 - Adiabatic invariants
- Basic collective plasma processes
 - Plasma frequency, skin depth (inertial length)
 - Debye length
- Collisional plasma: ideal MHD
 - MHD waves
 - Reduced MHD
- Weakly collisional plasma: non-ideal MHD
 - Ambipolar diffusion
 - Hall effect
- Collisionless plasma
 - hybrid - kinetics
 - gyro - kinetics

Plasma = collection of charged particles (2)

$$m_p \frac{d\mathbf{v}_p}{dt} = q_p \left[\mathbf{E} + \frac{\mathbf{v}_p \times \mathbf{B}}{c} \right]$$

electric field

$$\frac{dx_p}{dt} = v_p$$

Distribution function = probability of finding particle at certain point in phase space (x, v)

$$f(x, v, t) = \frac{2^N}{2 \pi \sigma v}$$

$$\int f d^3v = n(x) \leftarrow \text{particle number density}$$

$$\int v f d^3v = n(x) u(x) \leftarrow \text{particle momentum}$$

For a single particle,

$$f_1(x, v, t) = \delta(x - x_p) \cdot \delta(v - v_p)$$

For N particles

$$f_N(x, v, t) = \sum_{p=1}^N \delta(x - x_p) \delta(v - v_p)$$

How does f evolve with time?

$$\begin{aligned} \frac{\partial f_N}{\partial t} &= \frac{\partial}{\partial t} \sum_{p=1}^N \delta(x - x_p(t)) \delta(v - v_p(t)) = \\ &= \sum_{p=1}^N \left[+ \frac{dx_p}{dt} \frac{d}{dx_p} \delta(x - x_p) \delta(v - v_p) + \frac{dv_p}{dt} \frac{d}{dv_p} \delta(x - x_p) \delta(v - v_p) \right] = \\ &= - \sum_{p=1}^N \left\{ v_p \frac{\partial}{\partial x_p} \delta(x - x_p) \delta(v - v_p) + \frac{q_p}{m_p} \left(\mathbf{E} + \frac{\mathbf{v}_p \times \mathbf{B}}{c} \right) \frac{\partial}{\partial v_p} \delta(x - x_p) \delta(v - v_p) \right\} \end{aligned}$$

$\delta(v-v_p)$ is non-zero only for $v=v_p \Rightarrow$

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$$\Rightarrow v_p \frac{\partial}{\partial x} \delta \rightarrow v \frac{\partial}{\partial x} \delta$$

$$\frac{\partial f}{\partial t} = -v \frac{\partial}{\partial x} \left(\sum_p \delta(x-x_p) \delta(v-v_p) \right) - \frac{e_p}{m_p} \left(E + \frac{v \times B}{c} \right) \frac{\partial}{\partial v} \left(\frac{\delta(\cdot)}{\delta(\cdot)} \right)$$

$$= -v \cdot \nabla f - \frac{e}{m} \left(E + \frac{v \times B}{c} \right) \frac{\partial}{\partial v} f$$

↓

f evolves according to

(Klimontovich)

$$\frac{\partial f}{\partial t} + v \cdot \nabla f + \frac{e}{m} \left(E + \frac{v \times B}{c} \right) \frac{\partial f}{\partial v} = 0 \quad -\text{Kinetic equation}$$

E and B evolve according to Maxwell's equations

$$\nabla \cdot B = 0 \quad \nabla \cdot E = 4\pi \rho_c$$

$$\nabla \times B = \frac{4\pi}{c} J_c + \frac{1}{c} \frac{\partial E}{\partial t} \quad \underbrace{\nabla \times E = -\frac{1}{c} \frac{\partial B}{\partial t}}_{\text{Induction}}$$

charge density
 $\rho_c \equiv \sum_s q_s n_s$ ← sum over species
 current density
 $J_c \equiv \sum_s q_s n_s u_s$

E and B in Klimontovich equation are "microscopic"
 We typically want to have a smooth distribution function
 instead of delta functions $\Rightarrow f \rightarrow \langle f \rangle$, where $\langle \rangle$ is
 averaging over a small volume of phase space

$$E = \langle E \rangle + \delta E \quad B = \langle B \rangle + \delta B$$

$$\left(\frac{\partial}{\partial t} + v \cdot \nabla + \frac{e}{m} \left(\langle E \rangle + \frac{v \times \langle B \rangle}{c} \right) \cdot \frac{\partial}{\partial v} \right) f = - \underbrace{\left(\frac{e}{m} \left(\delta E + \frac{v \times \delta B}{c} \right) \frac{\partial f}{\partial v} \right)}_{\text{collision integral}}$$

Particle in uniform E and B fields

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$$\frac{dx}{dt} = v \quad m \frac{dv}{dt} = qE + q \frac{v \times B}{c}$$

$$1) B=0 \rightarrow v = \frac{q}{m} Et \quad - \text{linear acceleration}$$

$$2) E=0 \rightarrow \frac{dv}{dt} = \frac{q}{mc} v \times B$$

$$B = B_0 \hat{e}_z$$

$$\begin{aligned} \frac{dv_x}{dt} &= \frac{qB_0}{mc} v_y & v_y &= R v_y \\ \frac{dv_y}{dt} &= -\Omega v_x & \} & \Rightarrow v = v_0 e^{i\Omega t} \\ \Omega &= \frac{qB}{mc} = \text{Larmor frequency,} \\ & \text{gyro frequency} \end{aligned}$$



$$x = x_0 e^{-i\Omega t + i\pi/2}$$

$$x_0 = v_0/R$$

$$\Omega = \frac{qB}{mc} = \text{Larmor radius, gyro-radius.}$$

$$\rho_L = \frac{v}{\Omega} = \text{Larmor radius, gyro-radius.}$$

$$3) E, B \neq 0$$

$$\frac{dv}{dt} = \frac{q}{m} \left(E + \frac{v \times B}{c} \right)$$

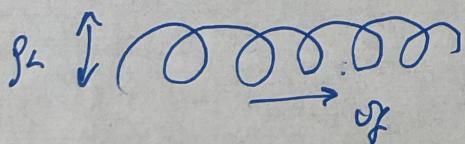
if $E < B$, there is a frame in which

$$E' = E + \frac{qv \times B}{c} = 0$$

$$v_f = c \frac{E \times B}{B^2}$$

in this frame $E' = 0, B' \neq 0 \rightarrow$ particle rotates around B

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motion could be described as superposition
of rotation and uniform motion



Invariants of particle motion

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Hamiltonian of a particle in electromagnetic field

$$H = \frac{1}{2m} \left(p - eA \right)^2 + e\varphi$$

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$$\nabla \times A = B$$

$$E = -\nabla \varphi - \frac{eA}{c}$$

Poincaré invariant (see, e.g. Goldstein 1980)

$$P = \oint p \cdot \frac{dq}{dx} dx \approx \text{const}$$

$$\frac{dp}{dt} = -\frac{\partial H}{\partial q} \quad \frac{dq}{dt} = \frac{\partial H}{\partial p}$$

p, q - canonical coordinates and momenta

$$\begin{aligned} \frac{dP}{dt} &= \oint \frac{d}{dt} \left(p \frac{dq}{dx} \right) dx = \oint \left\{ -\frac{\partial H}{\partial q} \frac{dq}{dx} + p \frac{d}{dx} \left(\frac{\partial H}{\partial p} \right) \right\} dx = \\ &= \underbrace{\oint \left\{ -\frac{\partial H}{\partial q} \frac{dq}{dx} - \frac{\partial H}{\partial p} \frac{dp}{dx} \right\} dx}_{-\frac{dH}{dx} dx \text{ integrates to zero}} + p \underbrace{\frac{\partial H}{\partial p}}_{0'' \text{ because the orbit is closed}} \Big|_{x_0}^{\lambda_0} \end{aligned}$$

$$\frac{dp}{dt} = 0$$

For an ion in uniform field,

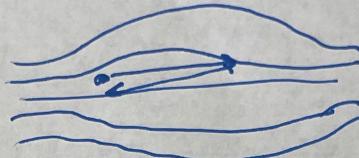
$$p = m(v - eA/mc) \rightarrow$$

$$\begin{aligned} \rightarrow \oint p \frac{dx}{dt} dt &= \int m v dx - \frac{e}{c} \int A dx = 2\pi m V_\perp \varphi - \\ &- \frac{e}{c} B (\pi r^2) \propto \frac{V_\perp^2}{B} \quad \mu = \frac{mr^2}{B} = \text{const} \\ &\text{"magnetic moment"} \end{aligned}$$

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Second adiabatic invariant

$$J = \oint p_{\perp} dl$$



$$\text{BETTER } \varepsilon = \frac{mv_x^2}{2} + \frac{mv_z^2}{2} = \text{const}$$

$$\mu = \frac{mv_z^2}{2eB} = \text{const}$$

$$\downarrow \\ v_z^2 = \frac{2eB}{m} \mu$$

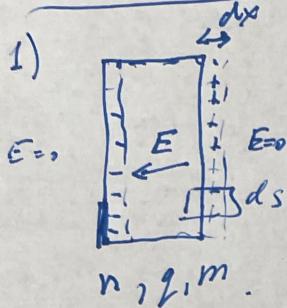
$$v_x^2 = \frac{2\varepsilon}{m} - \frac{2eB}{m} \mu = \frac{2}{m} (\varepsilon - eB\mu)$$

$$J = \oint dl \sqrt{\frac{2}{m} (\varepsilon - eB\mu)} \quad B = B(l)$$

$$v_{\parallel} = 0 \quad \text{when} \quad B(l) = \frac{\varepsilon}{e\mu}$$

Basic collective plasma processes

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$$\text{charge per unit area} \\ \sigma = \int n \cdot dx$$

$$E_0 = \frac{\sigma}{\epsilon_0} \Rightarrow E_0 = \frac{4\pi \sigma}{\epsilon_0}$$

$$E dS = \sigma \epsilon_0 dS \Rightarrow E = \frac{\sigma}{\epsilon_0}$$

restoring force per unit area

$$df = \sigma dS \cdot E = \frac{4\pi \sigma^2}{\epsilon_0} dS$$

$$dm = \text{mass per unit area} = m n dx$$

$$dm \ddot{x} = - \frac{4\pi \sigma^2}{\epsilon_0} dS = - \frac{4\pi \times^2 n^2 e^2}{\epsilon_0} dS$$

$$ds m n \ddot{x} = - \frac{4\pi \times^2 n^2 e^2}{\epsilon_0} dS$$

$$\ddot{x} = -x \cdot \frac{4\pi n^2 e^2}{m} = -\omega_p^2 x \quad \begin{matrix} \text{plasma oscillations} \\ \text{with frequency} \end{matrix}$$

$$\omega_p = \sqrt{\frac{4\pi n e^2}{m}} \quad \begin{matrix} \text{plasma} \\ \text{frequency} \end{matrix}$$

"Langmuir waves"

2) Place a charge in quasi-neutral plasma

charged particles form a Boltzmann distribution

$$n_e \approx n_0 e^{-\frac{e\gamma}{T}} \approx n_0 \left(1 - \frac{e\gamma}{T}\right)$$

$$n_h \approx n_0 e^{\frac{e\gamma}{T}} \approx n_0 \left(1 + \frac{e\gamma}{T}\right)$$

$$\beta_c = e n_e - e n_h \approx -2 n_0 \frac{e^2 \gamma}{T}$$

$$\nabla \cdot E = -\nabla^2 \gamma = 4\pi \beta_c = 4\pi Q \delta(r) - \underbrace{\sum_S 4\pi n_S \frac{q_S^2}{T_S}}_{\text{for arbitrary number of species}} \cdot \gamma$$

$$\lambda_D^{-2} = \sum_S 4\pi n_S \frac{q_S^2}{T_S} = \sum_S \lambda_{DS}^{-2}$$

$$\lambda_{DS} = \frac{T_S/m_S}{\omega_p} = V_{A,S}/\omega_p$$

$$-\nabla^2 \psi = 4\pi Q \delta(r) - \lambda_D^{-2} \psi \quad (8)$$

$$\psi = \frac{Q}{r} f(r) = \psi_0 + \frac{Q}{r} (f(r) - 1)$$

$$-\nabla^2 \psi = -\frac{1}{r^2} \frac{\partial^2}{\partial r^2} \left(r^2 \frac{\partial \psi_0}{\partial r} \right) - \frac{1}{r^2} \frac{\partial^2}{\partial r^2} \left(r^2 \frac{\partial^2}{\partial r^2} \left(\frac{Q}{r} (f-1) \right) \right)$$

$$\frac{1}{r^2} \frac{\partial^2}{\partial r^2} \left[r^2 \left\{ -\frac{Q}{r^2} (f-1) + \frac{Q}{r} f' \right\} \right] = \lambda_D^{-2} \frac{Q}{r} f$$

~~$$\frac{1}{r^2} \frac{\partial^2}{\partial r^2} \left[-f + 1 + r f' \right] = \frac{r^2}{\lambda_D^2} \frac{Q}{r} f$$~~

$$-f' + f' + r f'' = \frac{r^2}{\lambda_D^2} \frac{1}{r} f$$

$$f'' = \frac{1}{\lambda_D^2} f \Rightarrow f = e^{-r/\lambda_D}$$

$$\psi = \frac{Q}{r} \cdot e^{-r/\lambda_D}$$

Screening of charge at characteristic scale λ_D

if $r \gg \lambda_D$ plasma is quasi-neutral

Scales on plasma

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Larmor radius	$\rho_s = v/\omega_s$	$\mathcal{R}_s = \frac{\rho_s L}{\mu_s c}$
Skin-depth	$d_s = \sqrt{\omega_s \mu_s / \pi}$	$w_{ps} = \sqrt{\frac{4\pi n_s q_s^2}{m_s}}$
Debye length	$\lambda_{D_s} = v_{ms}/\omega_s$	$v_{ms} = \sqrt{T_s/m_s}$

λ_{mfp} - mean-free-path

L - typical scale

Dimensionless parameters

$\frac{L}{\lambda_{mfp}}$ - could be large or small

$$\frac{\rho_s}{d_s} = \frac{v_{ms} \lambda_{D_s}}{c / \omega_p} = \sqrt{\frac{T}{m}} \left(\frac{qB}{mc^2} \right) \sqrt{\frac{4\pi n q^2}{m}} = \sqrt{\frac{4\pi T n}{B^2}} = \sqrt{\frac{1}{2} \left(\frac{8\pi n T}{B^2} \right)}$$

$$\frac{\rho_s}{d_s} = \sqrt{\frac{B_s}{k}} \text{ plasma } \beta \quad \beta = \frac{\text{therm. magnetic pressure}}{\text{magneti. pres.}}$$

$\frac{L}{\rho_s}$ - typically $\gg 1 \Rightarrow$ plasma is magnetized

$\frac{\lambda_{D_s}}{d_s} = \frac{v_{ms}}{c}$ - small in non-relativistic plasmas \Rightarrow plasma is quasi-neutral

$$\frac{\partial f}{\partial t} + \mathbf{v} \cdot \nabla f + \frac{e}{m} \left(\mathbf{E} + \frac{\mathbf{v} \times \mathbf{B}}{c} \right) \frac{\partial f}{\partial \mathbf{v}} = C[f] \quad (1)$$

$\sum_s \int d\mathbf{v} C[f] = 0 \Rightarrow$ total number of particles is conserved during collision

$\sum_s \int d\mathbf{v} \mathbf{v} \cdot C[f] = 0 \Rightarrow$ total momentum is also conserved

$\sum_s \int d\mathbf{v} \frac{v^2}{2} C[f] = 0 \Rightarrow$ total energy is conserved

$$2\frac{\partial f}{\partial t} + \nabla_{\mathbf{x}} \cdot (\mathbf{v} f) + \nabla_{\mathbf{v}} \cdot \left(\frac{e}{m} \left(\mathbf{E} + \frac{\mathbf{v} \times \mathbf{B}}{c} \right) f \right) = C[f]$$

$$\int d\mathbf{v} \dots$$

$$\frac{\partial n}{\partial t} + \nabla_{\mathbf{x}} \cdot (n \mathbf{u}) = 0 \quad - \text{continuity equation}$$

$$\int \sigma d\mathbf{v} \dots$$

~~$$\frac{\partial n u}{\partial t} + (\mathbf{u} \cdot \nabla_{\mathbf{x}})(n u) + \frac{e}{m} \left(\mathbf{E} + \frac{\mathbf{v} \times \mathbf{B}}{c} \right) = 0$$~~

$$\begin{aligned} & \int \cancel{\sigma} \cdot \nabla_{\mathbf{x}} \left(f \frac{\mathbf{v} \times \mathbf{B}}{c} \right) d\mathbf{v} = \\ & = - \cancel{\int f \frac{\mathbf{v} \times \mathbf{B}}{c} d\mathbf{v}}$$

pressure tensor

$$\int \sigma \frac{\partial f}{\partial t} d\mathbf{v} = \frac{\partial}{\partial t} (n u)$$

$$\int \sigma \cdot \nabla_{\mathbf{x}} (f \mathbf{v}) d\mathbf{v} = \nabla_{\mathbf{x}} \cdot \int f \mathbf{v} \sigma d\mathbf{v} = \nabla_{\mathbf{x}} \left(g u u + \underbrace{P}_{\text{pressure}} \right)$$

$$\int \sigma \cdot \nabla_{\mathbf{x}} (E f) d\mathbf{v} = - E n u$$

$$\int \sigma \cdot \nabla_{\mathbf{x}} \left(\frac{\mathbf{v} \times \mathbf{B}}{c} f \right) d\mathbf{v} = - \frac{n u \times \mathbf{B}}{c}$$

Sum over species

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$$\sum_k n_k E = \rho c E = 0 \quad \text{for } L \gg \lambda_D$$

$$\sum_k n_k \frac{u_{kx} v_D}{c} = \frac{Jx\delta}{c}$$

↙

$$\frac{\partial h u}{\partial t} + \nabla \cdot (\rho u u + P) = \frac{Jx\delta}{c}$$

$\rho u \cdot \nabla u$

$P I + \Pi \leftarrow \text{various stress}$

$$u = \frac{1}{\rho_m} \sum_m m_m u_m$$

center of mass
velocity

Ideal hydrodynamics

$$\frac{\partial \phi}{\partial t} + \nabla \cdot (\beta u) = 0$$

$$\frac{\partial \phi u}{\partial t} + \rho u \cdot \nabla u = - \nabla P$$

1 characteristic velocity - sound speed

$$C_s^2 = \frac{P}{\rho}$$

$\delta u \approx C_s$ - strong compression, shock waves

$\delta u \ll C_s$ - no shocks, sound waves
equilibrate pressure

Ideal MHD

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$$\frac{\partial \mathbf{B}}{\partial t} + \nabla \cdot (\mathbf{B} \mathbf{u}) = 0$$

$$\frac{\partial \mathbf{B}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{B} = -\nabla P + \frac{\mathbf{J} \times \mathbf{B}}{c}$$

$$2 \quad \nabla \cdot \mathbf{B} = \frac{\mu_0}{c} J \quad \leftarrow \text{ neglect displacement current}$$

$$1 \quad \frac{1}{c} \frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E} \quad \leftarrow \mathbf{E} + \frac{\mathbf{u} \times \mathbf{B}}{c} = 0 \quad \text{"frozen-in condition"}$$

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{u} \times \mathbf{B})$$

new characteristic velocity associated with $\mathbf{J} \times \mathbf{B}$ for

$$v_A \approx \frac{B_0}{\sqrt{\mu_0 \rho}} \quad \text{Alfvén speed}$$

1 continuity equation

3 momentum equations

3 induction equations + 1 constant $\mathbf{e} \cdot \mathbf{B} = 0$

1 energy equation

7 wave solution

1 entropy wave

2 Alfvén waves

2 slow magnetosonic modes

2 fast magnetosonic modes

} analogous to sound waves.

Non-ideal MHD

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$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{u} \times \mathbf{B})$$

what is \mathbf{u} ?

Consider a collisional plasma

$$m_n n_n (\partial_t + u_n \nabla) u_n = -\nabla p_n + R_{ni} + R_{ne}$$

$$m_i n_i (\partial_t + u_i \nabla) u_i = -\nabla p_i + R_{in} + R_{ei} + q_0 n_e (E + \frac{u_i \times B}{c})$$

$$m_e n_e (\partial_t + u_e \nabla) u_e = -\nabla p_e + R_{en} + R_{ei} - e n_e (E + \frac{u_e \times B}{c})$$

$$\begin{aligned} \frac{\partial \mathbf{B}}{\partial t} &= \nabla \times [(u_e - u_n) \times \mathbf{B} + \\ &\quad + (u_e - u_i) \times \mathbf{B} + \text{---> Hall effect} \\ &\quad + (u_i - u_n) \times \mathbf{B} + \text{---> Ambipolar diffusion} \\ &\quad + \underbrace{u_n \times \mathbf{B}}_{\approx u \times \mathbf{B}}] \\ &\quad \text{center of mass} \end{aligned}$$

$$\frac{\partial u_n}{u \times \mathbf{B}} \sim \frac{1}{R_m} = \frac{y}{V_A l_B} \sim \left(\frac{d_e}{l_B} \right) \left(\frac{d_e}{V_A T_{en}} \right) \begin{array}{l} \swarrow \text{small} \\ \curvearrowleft \end{array} \begin{array}{l} \curvearrowright \text{could be} \\ \curvearrowright \text{large} \end{array}$$

$$\frac{\text{Hall}}{u \times \mathbf{B}} \sim \frac{|V_{Ae}|}{|u_n|} \sim \left(\frac{d_i}{l_0} \right) \left(\frac{S_m}{S_{ai}} \right)^{Y_2} \left(\frac{V_A}{u_n} \right) \alpha_T$$

$\ll 1 \quad \sim 1$

$$\frac{\text{Antipolar}}{n_0 B} \sim \left| \frac{R_{\text{ext}} v_n}{g_n u_n} \right| \sim \left(\frac{j \times B}{c} \right) \frac{e n_i}{g_n u_n} \left(\frac{v_A}{u_n} \right)$$

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Collisionless plasma

fully - kinetic

$$\frac{\partial f_i}{\partial t} + \mathbf{v} \cdot \nabla f_i + \frac{q_i}{m_i} \left(\mathbf{E} + \frac{\mathbf{v} \times \mathbf{B}}{c} \right) \frac{\partial f_i}{\partial \mathbf{v}} = 0$$

$$\frac{\partial f_e}{\partial t} + \mathbf{v} \cdot \nabla f_e + \frac{q_e}{m_e} \left(\mathbf{E} + \frac{\mathbf{v} \times \mathbf{B}}{c} \right) \frac{\partial f_e}{\partial \mathbf{v}} = 0$$

hybrid - kinetic : $\ell \sim d_i, g_i \gg d_e, g_e$
 $m_e/m_i \ll 1$

Equation of motion for electrons gives E

$$m_e \frac{d\mathbf{v}_e}{dt} = \mathbf{F} + \frac{q_e \mathbf{v}_e \times \mathbf{B}}{c} - \nabla \phi_e$$

MHD waves

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$$\frac{\partial \mathbf{g}^u}{\partial t} = -\nabla \cdot (\mathbf{g} u u + p \mathbf{I} + \frac{B^2}{8\pi} \mathbf{I} - \frac{B^2}{4\pi} \mathbf{b} \mathbf{b})$$

$$\frac{\mathbf{j} \times \mathbf{B}}{c} = \frac{(\mathbf{g} \times \mathbf{B}) \times \mathbf{B}}{4\pi}$$

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{u} \times \mathbf{B})$$

$$\frac{\partial \mathbf{g}}{\partial t} + \nabla \cdot (\mathbf{g} u) = 0$$

$$K \cdot \delta \mathbf{B} = 0 \Rightarrow \delta \mathbf{B}_u = 0$$

$$\frac{\partial e}{\partial t} + \nabla \cdot (e u) = -p \nabla \cdot u \rightarrow \left(\frac{\partial}{\partial t} + \mathbf{u} \cdot \nabla \right) s = 0$$

$$s = (\gamma - 1)^{-1} \ln [P \mathbf{g}^{-\gamma}]$$

Linearization

$$\mathbf{B} = \mathbf{B}_0 + \delta \mathbf{B}$$

$$u = u_0 + \delta u, \quad u_0 = 0$$

$$\mathbf{g} = \mathbf{g}_0 + \delta \mathbf{g}$$

$$p = p_0 + \delta p$$

$$\delta g = g_0 \frac{K \cdot \delta u}{\omega}$$

$$-i\omega \delta g + i g_0 K \cdot \delta u = 0$$

$$i\omega [\delta p \mathbf{g}^{\gamma} - \gamma \delta g p \mathbf{g}^{\gamma-1}] = 0 \rightarrow \omega = 0 \text{ entropy wave}$$

$$\frac{\delta p}{p} = \gamma \frac{\delta g}{g}$$

$$-i\omega g_0 \delta u = -iK \delta p + i(k \times \delta \mathbf{B}) \frac{x B_0}{4\pi}$$

$$-i\omega \delta B = ik \times (\delta u \times B_0)$$

$$-i\omega \delta u = -K \underbrace{\gamma \frac{p_0}{g_0^2} g_0}_{C_s^2} \frac{K \cdot \delta u}{\omega} + \frac{[K \times \delta B] \times B_0}{4\pi g_0}$$

$$-i\omega \delta u = -K C_s^2 \frac{K \cdot \delta u}{\omega} + \frac{[k \times \delta B] \times B_0}{4\pi g_0}$$

$$K = K_H$$

if $\delta u = \delta u_H$

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$$-i\omega \delta B = iK_H \times (\delta u \times B_0) = 0 \Rightarrow \delta B \approx 0$$

$$-\omega \delta u = -K \frac{B_0}{\rho_0} S \cdot \frac{K_H \delta u_H}{\omega}$$

\swarrow
Sound waves $\omega = k_H c_s$

$$\delta u = \delta u_H$$

$$-i\omega \delta B_\perp = i\delta u_\perp K_{H\perp} B_0$$

$$\delta B_\perp = \delta u_\perp \frac{k_H}{\omega} \frac{B_0}{\rho_0}$$

$$-\omega \delta u_\perp = \frac{(K_H \times \delta u_H \frac{k_H B_0}{\omega \rho_0}) \times B_0}{4\pi \rho_0} =$$

$$= -\frac{K_H^2 V_A^2}{\omega} \delta u_H$$

$$V_A = \sqrt{\frac{B_0}{4\pi \rho_0}}$$

Alfvén waves

$$K = K_H + K_\perp$$

$$(\omega^2 - K_{H\perp}^2 V_A^2) \left[\omega^2 - K_H^2 V_A^2 - K_\perp^2 \frac{\omega^2}{\omega^2 + c_S^2} \right] = 0$$

$$\omega^2 = \frac{K^2 (c_S^2 + V_A^2)}{2} \pm \sqrt{\frac{K^2 (c_S^2 + V_A^2)^2}{4} - K_H^2 V_A^2 K^2 c_S^2}$$

fast and slow modes

Mean free path

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$$n \sigma \lambda_{\text{mfp}} = 1$$

$$\sigma \sim \pi b^2 \quad \frac{\pi^2}{b} \sim T \rightarrow \lambda_{\text{mfp}} \propto \frac{T^{1/2}}{n} T^2 / n$$

$$\tau_{ei} \sim \frac{3\sqrt{m_e e^3 k_B}}{4\sqrt{2\pi m_i n_e} N_e e^4}$$

$$\tau_{ii} \sim \frac{3}{4\sqrt{\pi}} \frac{m_e T_e^{3/2}}{m_i N_i e^4}$$

$$\lambda_{\text{mfp},e} = \frac{3}{4\sqrt{\pi}} \frac{T_e^{1/2}}{N_e} \frac{1}{N_e e^4}$$

$$\lambda_{\text{mfp},i} \approx \frac{3\sqrt{2}}{4\sqrt{\pi}} \frac{T_i^{1/2}}{N_i N_i e^4} \approx \lambda_{\text{mfp},e}$$

$$\lambda_{\text{mfp},i} \approx \lambda_{\text{mfp},e} = \lambda_{\text{mfp}}$$